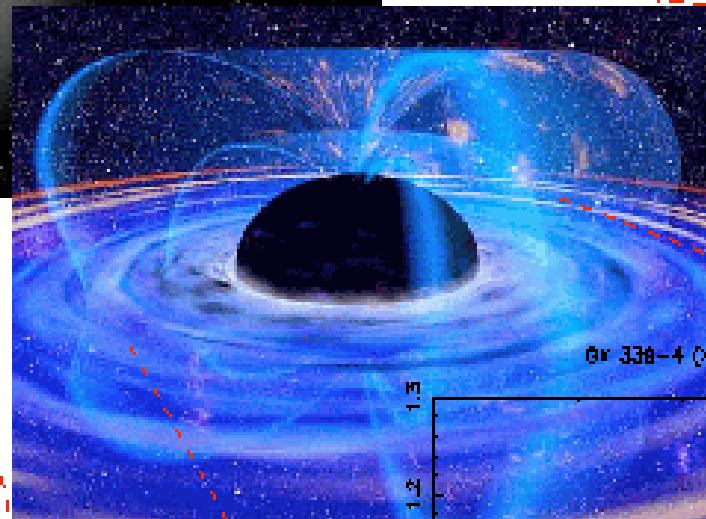
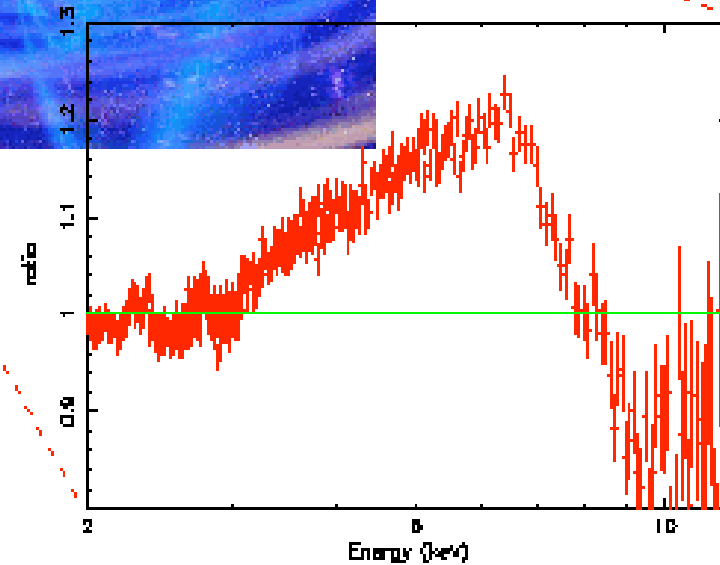


# Revealing Relativity: Galactic Black Holes in the Con-X Era



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(on behalf of a larger group)



# Old Con-X Objectives for X-ray Binaries

- “Use X-ray radial velocity measurements to determine the mass function of black holes, neutron stars, and white dwarfs in binary systems.”

— *For the vast majority of systems, this will be done with narrow Bowen fluorescence lines from the companion.*

- “Measure the gravitational redshift at the surface of white dwarfs in magnetic CVs using the Fe K fluorescence lines.”

— *Resolution of  $\sim 3000$  is needed.*

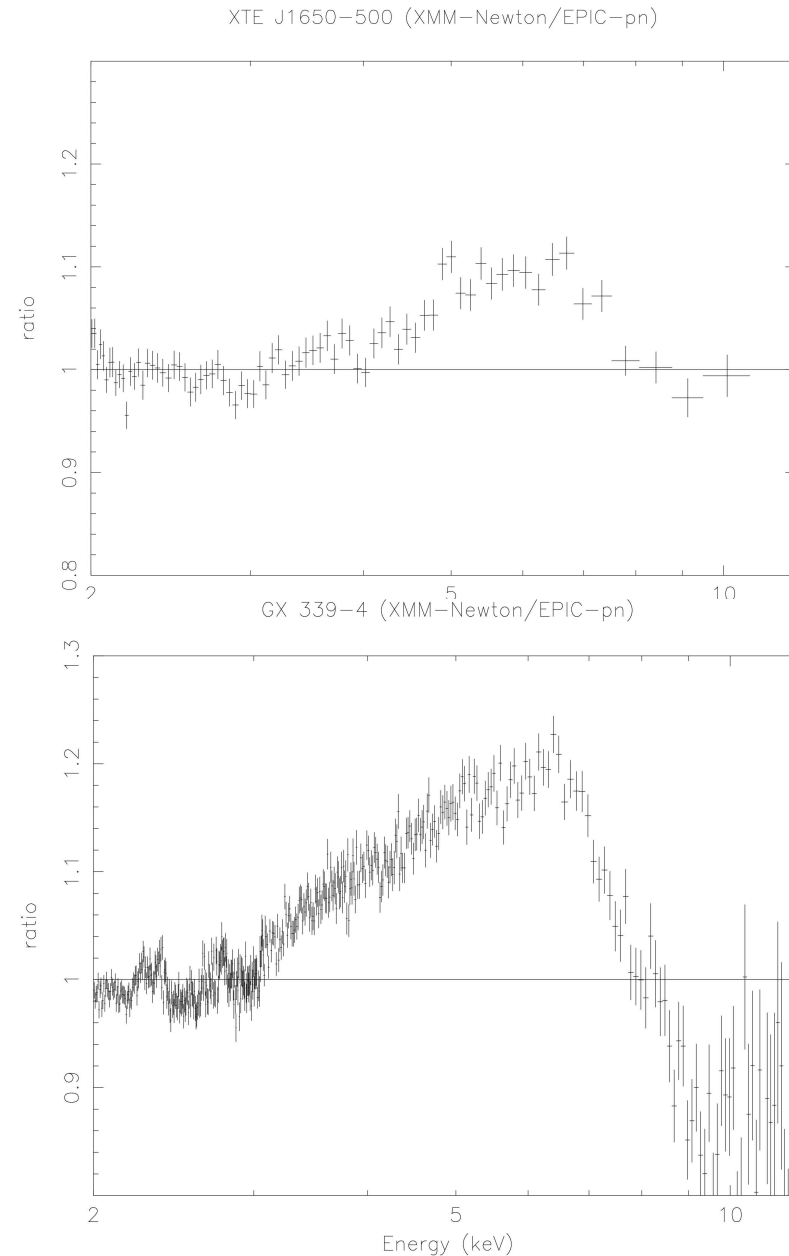
# New Objectives for Galactic Black Holes

- Search for black hole spin, range of spin parameters via skewed Fe K lines. Search for other skewed lines.
- Probe the corona-disk interaction: quasi-spherical corona, pancake, or magnetic flaring loops?
- Determine the nature of low-luminosity flows: ADAF, jet, or ADAF+jet, and where is the disk?

# Relativistic Fe K Lines

- Rule-out 6  $R_{\text{grav}}$ .  
\_  $R = 1.24 R_{\text{grav}}$ .
- Rule-out  $q = 3$  ,  $J(r) = r^{(-q)}$   
\_  $q = 5.5-6$
- Simultaneous [Chandra/XMM-N] and RXTE for continuum.  
\_ Lines are largely independent of the continuum model chosen.

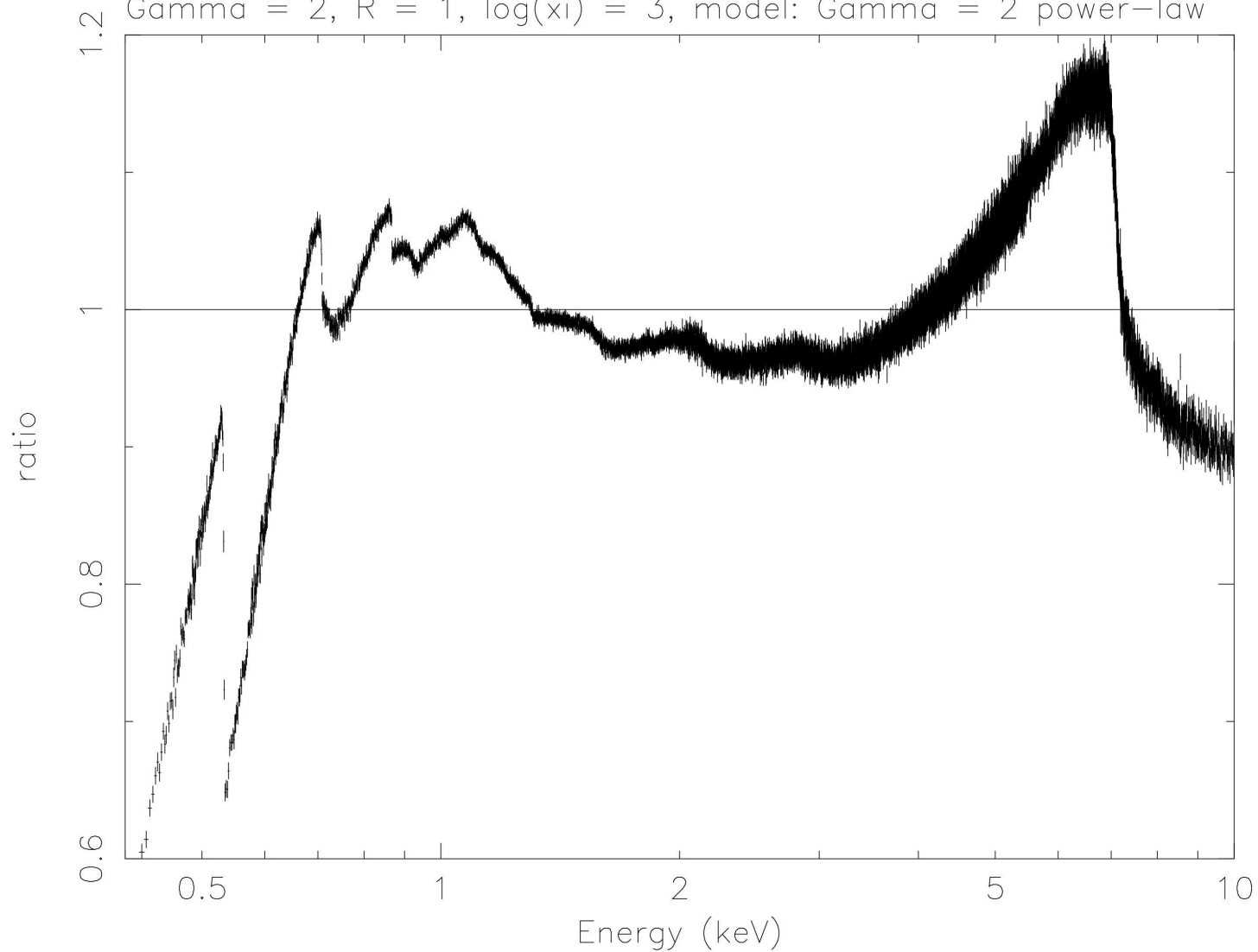
*We need to be able to handle ~1 Crab sources.*



# Multiple Relativistic Lines

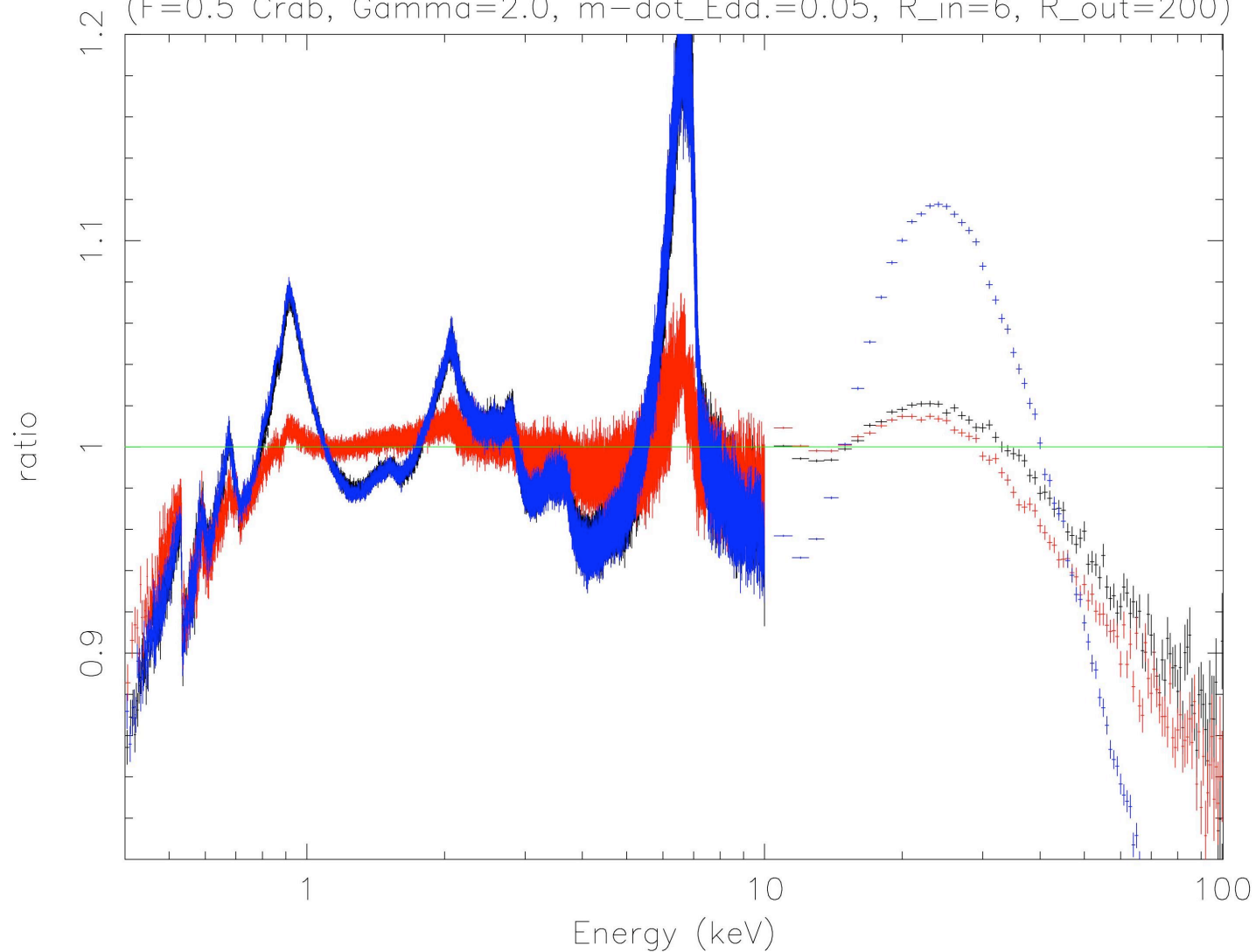
Con-X Calorimeter: 1 Crab (Kerr BH) 100 ksec Reflection Spectrum

Gamma = 2, R = 1, log(xi) = 3, model: Gamma = 2 power-law

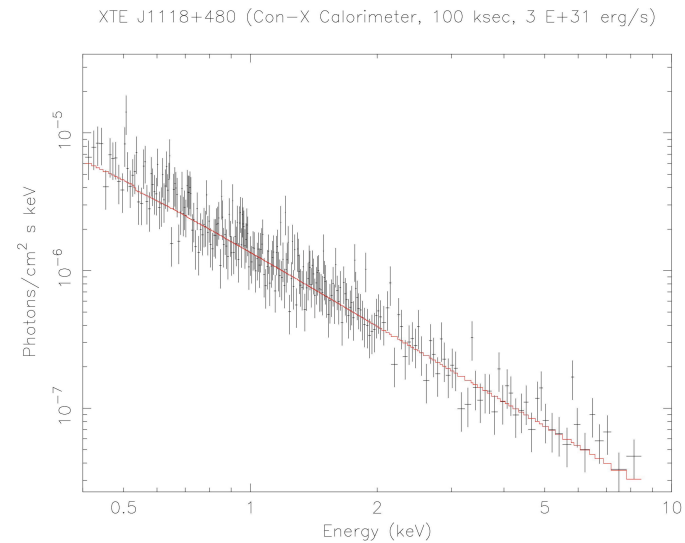
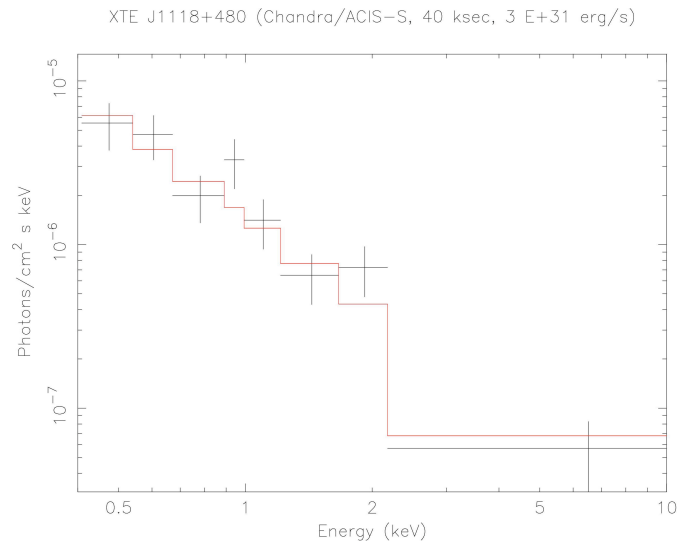
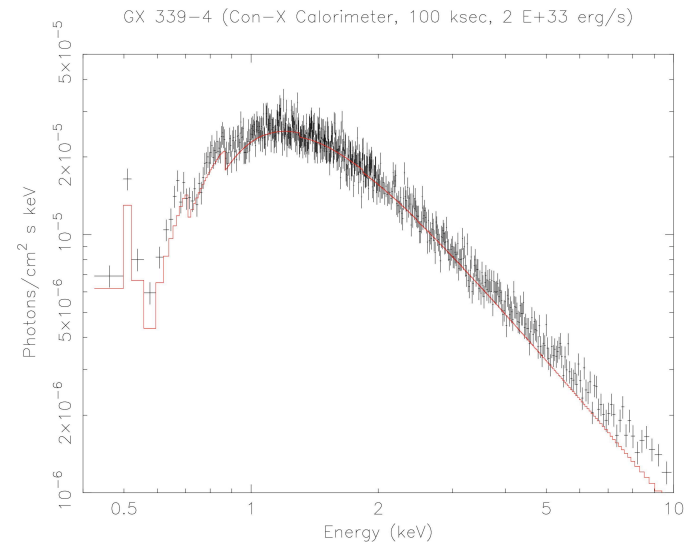
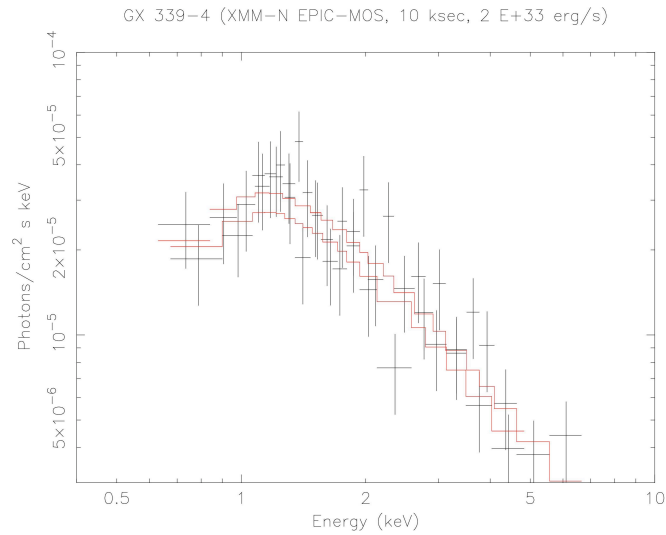


# The Disk-Corona Interaction

XION Reflection Models, Black = lamp-post, red=sphere, blue=flares  
( $F=0.5$  Crab,  $\Gamma=2.0$ ,  $\dot{m}_{\text{Edd}}=0.05$ ,  $R_{\text{in}}=6$ ,  $R_{\text{out}}=200$ )



# Galactic Black Holes in Quiescence



# Con-X Will be Revolutionary for GBHs

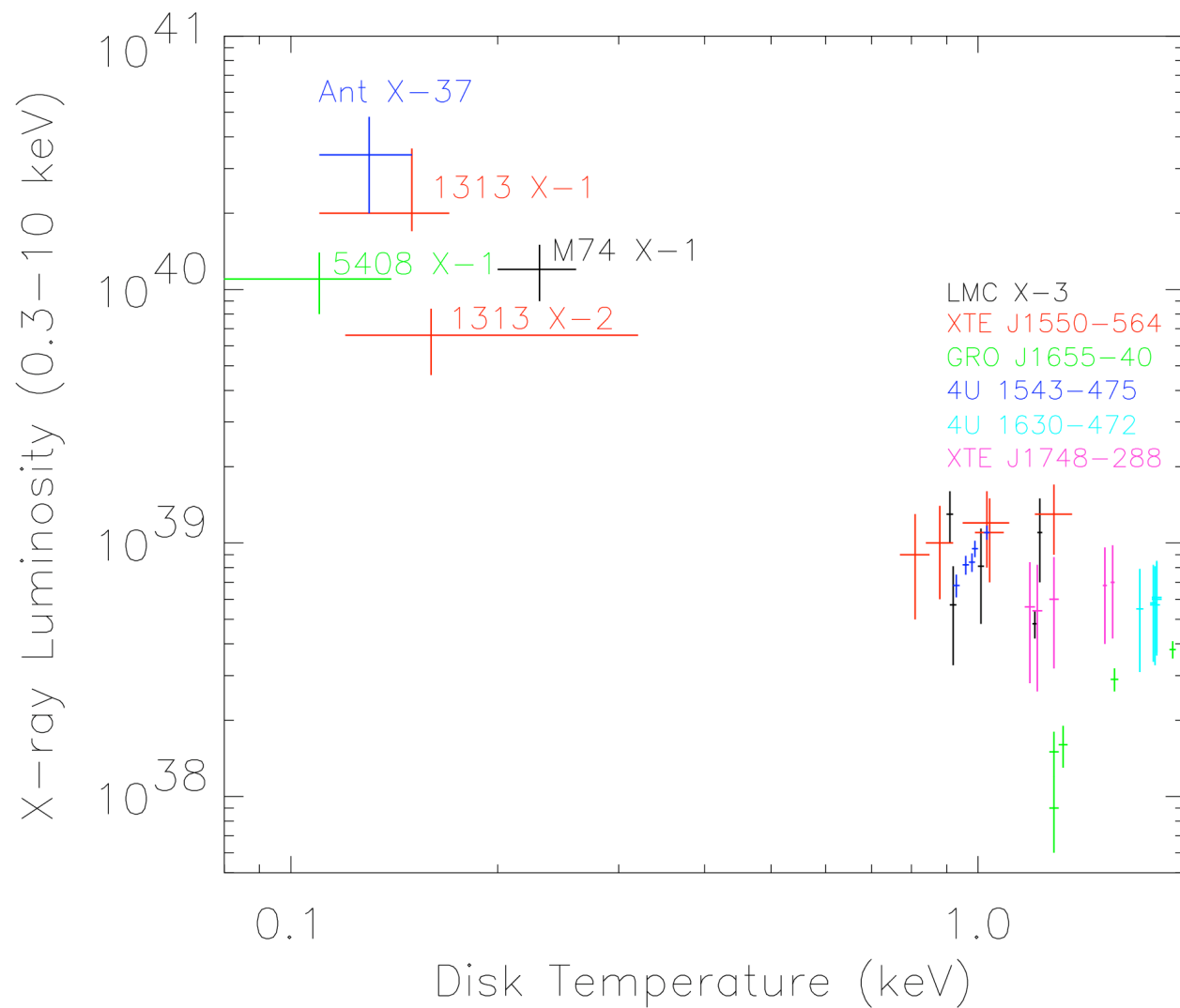
- Spin may be revealed with multiple relativistic lines.
- Even more exotic physics may be probed (B-Z effect).
- The jet ejection mechanism can be constrained by probing the coronal geometry, corona-disk interaction.
- Multiple-component spectra in quiescent systems.

*\_ We need to be able to handle 1 Crab sources, whether through calorimeter design, defocusing, or both.*

(won't someone please ask me about ULXs / IMBHs?)



# Con-X: The Machine for ULXs



# Con-X Will Reveal Reflection in ULXs

NGC 1313 X-1:  $L_X = 2 \times 10^{40}$   
Disk Temp: 150 eV

